


$$\begin{matrix} 103m \\ 45 \quad \text{Rh} \quad 58 \\ 58 \end{matrix}$$

### Evaluation of the decay data

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The initial evaluation was completed in May 2013 and then successively revised in January 2015, April 2021, April 2023 and June 2024. The present evaluation was carried out with a literature cut-off date of June 2024.

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## 1 Decay Scheme

$^{103m}$ Rh undergoes 100% IT decay by means of a highly-converted 39,753-keV gamma transition.  
*Le  $^{103m}$ Rh se désexcite par une transition gamma de 39,753 keV fortement convertie.*

## 2 Nuclear Data

$$\begin{aligned} T_{1/2}(^{103m}\text{Rh}) &: 56,115 \quad (16) \quad \text{min} \\ Q^{IT}(^{103m}\text{Rh}) &: 39,753 \quad (6) \quad \text{keV} \end{aligned}$$

### 2.1 Gamma Transitions and Internal Conversion Coefficients

	Energy (keV)	P <sub>γ+ce</sub> (%)	Multipolarity	α <sub>K</sub>	α <sub>L</sub>	α <sub>M</sub>	α <sub>T</sub>
γ <sub>1,0</sub> (Rh)	39,753 (6)	100	E3+M4	142 (7)	1051 (30)	214 (6)	1440 (40)

## 3 Atomic Data

### 3.1 Rh

$$\begin{aligned} \omega_K &: 0,809 \quad (4) \\ \bar{\omega}_L &: 0,0494 \quad (12) \\ n_{KL} &: 0,987 \quad (4) \end{aligned}$$

### 3.1.1 X Radiations

	Energy (keV)	Relative probability	
X <sub>K<sub>tot</sub></sub>	20,073 - 23,173	183,7	
K $\alpha_2$	20,073	52,83	K $\alpha$
K $\alpha_1$	20,215	100	
K $\beta_3$	22,699	8,50	K' $\beta_1$
K $\beta_1$	22,725	16,55	
KN	23,167 - 23,173	4,97	
X <sub>L<sub>tot</sub></sub>	2,378 - 3,067	96,8	
X <sub>M<sub>tot</sub></sub>	0,189 - 0,546	2,685	

### 3.1.2 Auger Electrons

	Energy (keV)	Relative probability
Auger K <sub>tot</sub>	16,281 - 22,838	143
KLL	16,281 - 17,083	100
KLX	19,133 - 20,146	39,2
KXY	21,922 - 22,838	3,69
Auger L <sub>tot</sub>	0,004 - 3,133	6373
Auger M <sub>tot</sub>	0,021 - 0,390	17540
Auger N <sub>tot</sub>	0,001 - 0,064	19330

## 4 Electron Emissions

		Energy (keV)	Electrons (per 100 disint.)
eAN <sub>tot</sub>	(Rh)	0,001 - 0,064	259 (5)
eAM <sub>tot</sub>	(Rh)	0,021 - 0,390	235 (5)
eAL <sub>tot</sub>	(Rh)	0,004 - 3,133	85,4 (19)
eAK <sub>tot</sub>	(Rh)	16,281 - 22,838	1,92 (9)
	KLL	16,281 - 17,083	1,34 (6)
	KLX	19,133 - 20,146	0,525 (24)
	KXY	21,922 - 22,838	0,0495 (22)
ec <sub>1,0</sub> T	(Rh)	16,533 - 39,751	99,9306 (15)
ec <sub>1,0</sub> K	(Rh)	16,533 (6)	9,85 (45)
ec <sub>1,0</sub> L	(Rh)	36,341 - 36,749	73 (3)
ec <sub>1,0</sub> M	(Rh)	39,126 - 39,446	14,9 (5)
ec <sub>1,0</sub> N	(Rh)	39,672 - 39,751	2,18 (8)

## 5 Photon Emissions

### 5.1 X-Ray Emissions

		Energy (keV)	Photons (per 100 disint.)
XM <sub>tot</sub>	(Rh)	0,189 - 0,546	0,1184 (23)
XL <sub>tot</sub>	(Rh)	2,378 - 3,067	4,27 (10)
XK <sub>tot</sub>	(Rh)	20,073 - 23,173	8,1 (4)
XK $\alpha_2$	(Rh)	20,073	2,33 (10)
XK $\alpha_1$	(Rh)	20,215	4,41 (20)
XK $\beta_3$	(Rh)	22,699	0,375 (17)
XK $\beta_1$	(Rh)	22,725	0,730 (33)
XKN	(Rh)	23,167 - 23,173	0,219 (10)

## 5.2 Gamma Emissions

Energy (keV)	Photons (per 100 disint.)
$\gamma_{1,0}(\text{Rh})$ 39,753 (6)	0,0694 (15)

## 6 Main Production Modes

$\left\{ \begin{array}{l} {}^{103}\text{Pd} \text{ EC decay} \\ \text{Possible impurities: } {}^{103}\text{Pd}, {}^{109}\text{Pd}, {}^{111}\text{Pd}, {}^{111m}\text{Pd} \end{array} \right.$

$\left\{ \begin{array}{l} {}^{103}\text{Ru} \beta^- \text{ decay} \\ \text{Possible impurities: } {}^{97}\text{Ru}, {}^{103}\text{Ru}, {}^{105}\text{Ru}, {}^{106}\text{Ru} \end{array} \right.$

$\left\{ \begin{array}{l} {}^{103}\text{Rh(n,n')} {}^{103m}\text{Rh} \\ \text{Possible impurity: } {}^{104m}\text{Rh} \end{array} \right.$

${}^{103}\text{Rh(d,d')} {}^{103m}\text{Rh}$

${}^{103}\text{Rh}(\alpha,\alpha') {}^{103m}\text{Rh}$

${}^{103}\text{Rh}(\gamma,\gamma') {}^{103m}\text{Rh}$

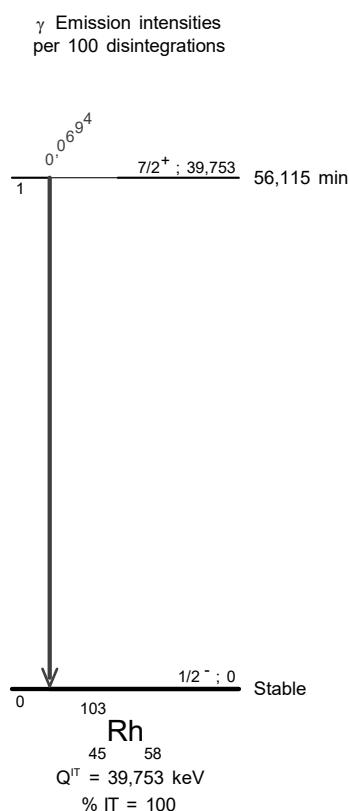
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**<sup>103m</sup>Rh – Comments on evaluation of decay data**  
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**<sup>103m</sup>Rh IT decay separated from <sup>103</sup>Pd (100% EC decay)**

**Decay Scheme**

<sup>103m</sup>Rh undergoes 100% IT decay ( $T_{1/2} = 56.115$  minutes) to the ground state of <sup>103</sup>Rh ( $Q_{IT} = 39.753$  (6) keV). A simple decay scheme can be proposed on the basis of the depopulation of the metastable state by means of the single 39.753-keV gamma ray to the ground state of <sup>103</sup>Rh.

**Nuclear Data**

There is a strong interest in the nuclear structure of <sup>103</sup>Rh and the IT decay of <sup>103m</sup>Rh because of the adoption of <sup>103m</sup>Rh as a neutron flux monitor.

**Half-life of <sup>103m</sup>Rh**

The measurements of 1967VuZZ, 1969KoZW, 1973Gu06, 1974Sa15, 1978La21 and 1981Va11 were adopted to give a weighted mean half-life of 56.115 (16) min based on LWM-NRM-Rajeval analyses (Avetools code, version 3.0, and V.AveLib 06-22 code, both ENSDF utility programs).

Reference	Half-life (min)	Selected detail, issues and uncertainties
1944Fl01	$48 \pm 5$	<sup>103</sup> Rh(n,n') reaction – outlier (Chauvenet criterion), and therefore not included in analytical procedures
1945Wi03/1945Wi12	$45 \pm 1$	<sup>103</sup> Rh(X, $\gamma$ ) reaction – outlier (Chauvenet criterion), and therefore not included in analytical procedures
1947Fl03	$52 \pm 2$	<sup>103</sup> Rh(d,d') reaction – outlier (Chauvenet criterion), and therefore not included in analytical procedures
1950Me26	56	separation of Pd target and <sup>103m</sup> Rh – uncertainty not specified, and therefore not included in analytical procedures
1957Jo19	$57.5 \pm 0.5$	<sup>103</sup> Rh(n,n') reaction – outlier (Chauvenet criterion), and therefore not included in analytical procedures
1967VuZZ	<b><u>56.6 ± 0.4</u></b>	<sup>103</sup> Rh(n,n') reaction; 4π proportional counter and thin NaI crystal
1969KoZW	<b><u>56.6 ± 0.6</u></b>	<sup>102</sup> Ru(n, $\gamma$ ) <sup>103</sup> Ru( $\beta^-$ ) <sup>103m</sup> Rh; radiochemical separation, NaI(Tl) scintillation counter
1972Pa10	57.5	<sup>103</sup> Rh(n <sub>14.7keV</sub> ,n') reaction, Si(Li) detector – uncertainty not specified, and therefore not included in analytical procedures
1973Gu06	<b><u>56.116 ± 0.009</u></b>	<sup>nat</sup> Pd(n,xn) <sup>103</sup> Pd(EC) <sup>103m</sup> Rh; three counting techniques: liquid scintillation counter, 2π-methane flow counter, and NaI(Tl) crystal – weighted mean of three sets of measurements is 56.116 (7), compared with a value of 56.116 (9) quoted by 1973Gu06; uncertainty adjusted to $\pm 0.016$ to reduce weighting to ~ 0.50
1974Sa15	<b><u>56.3 ± 0.6</u></b>	<sup>103</sup> Rh(n,n') reaction; decay curve followed over many half-lives, NaI(Tl) X-ray detector
1978La21	<b><u>56.112 ± 0.028</u></b>	ion exchange of <sup>103m</sup> Rh from <sup>103</sup> Pd solutions; decay curves involving 15 sources followed over three half-lives by means of an ill-defined photon spectrometer
1979VaZE	56.1 ± 0.1	preliminary measurement - replaced by half-life determined by 1981Va11
1981Va11	<b><u>56.114 ± 0.020</u></b>	ion exchange of <sup>103m</sup> Rh from <sup>103</sup> Pd solutions; decay curves followed over two and three half-lives, NaI(Tl) detector – weighted mean of two sets of measurements is 56.114 (12), compared with a value of 56.114 (20) quoted by 1981Va11
<b>Recommended value</b>	<b><u>56.115 ± 0.016</u></b>	recommended uncertainty adjusted from $\pm 0.011$ to $\pm 0.016$ , to align with smallest uncertainty of the values used to calculate weighted-mean value

Limitation of relative statistical weight method (LWM), normalized residual method (NRM), Rajeval technique, expected value method (EVM), bootstrap method and Mandel-Paule approach were considered in the analysis of the data set.

Analytical method	Half-life (min)	$\chi^2/(N-1)$	$\chi^2/(N-1)_{critical}$
UWM	$56.31 \pm 0.16$	–	–
LWM	$56.115 \pm 0.011$	0.45	2.21 for rejection at 95.0% confidence level
NRM	$56.115 \pm 0.011$	0.45	2.21 for rejection at 95.0% confidence level
Rajeval	$56.115 \pm 0.011$	0.45	2.21 for rejection at 95.0% confidence level
EVM	$56.13 \pm 0.08$	–	97.0% confidence level
Bootstrap	$56.22 \pm 0.17$	15.8	–
Mandel-Paule	$56.115 \pm 0.011$	0.45	–

Half-life of  $(56.115 \pm 0.016)$  minutes recommended for <sup>103m</sup>Rh, as quantified by LWM-NRM analytical procedures (1973Gu06 uncertainty increased to  $\pm 0.016$  in order to adjust weighting to  $\sim 0.50$ ).

### Q value

A Q<sub>IT</sub> value for <sup>103m</sup>Rh of 39.753 (6) keV has been adopted (2021Ko07).

### Gamma-ray energies and emission probabilities

#### Energies

The well-defined nuclear level energy for the first-excited metastable state was adopted from the Q<sub>IT</sub> value, along with the recommended energy and associated uncertainty of the gamma-ray emission to the ground state of <sup>103</sup>Rh (2021Ko07).

#### Adopted energies, spins and parities for the relevant nuclear levels of <sup>103</sup>Rh (see also 2009De29, 2000KoZT).

Nuclear level number	Nuclear level energy (keV)	Spin and parity	Half-life	
0	0.0	1/2 -	stable	<sup>103</sup> Rh
1	$39.753 \pm 0.006$	7/2 +	56.115 (16) min	<sup>103m</sup> Rh

#### Emission Probabilities

A simple decay scheme has been assembled for the IT decay of <sup>103m</sup>Rh. The 39.753-keV gamma transition depopulates the <sup>103m</sup>Rh nuclear level 100%, and is defined as (99.935%E3 + 0.065%M4) from spin and parity considerations ( $7/2^+ \rightarrow 1/2^-$ ), internal conversion coefficients, and shell and subshell ratios (1950Ko10, 1952Co16, 1955Av11, 1955Dr43, 1967Br04, 1967VuZZ, 1968Ma08, 1969Gr13, 1969Le17, 1970NiZV, 1970Pe04, 1972Br02, 1972Pa10, 1974CzZY, 1975Cz03, 1975Ma32, 1976Ma37, 1979VaZE, 1999Sa78, 2018Ni14). Recommended internal conversion coefficients have been determined by means of the BrIcc code (frozen orbital approximation) of Kibédi *et al.* (2008Ki07), based on the theoretical model of Band *et al.* (2002Ba85, 2002Ra45).

#### Gamma-ray emission from 100% IT decay of <sup>103m</sup>Rh: multipolarity and theoretical internal-conversion coefficients (frozen orbital approximation).

E <sub>γ</sub> (keV)	Multipolarity	ω <sub>K</sub>	ω <sub>L</sub>	ω <sub>L1</sub>	ω <sub>L2</sub>	ω <sub>L3</sub>	ω <sub>M</sub>	ω <sub>N</sub>	ω <sub>O</sub>	ω <sub>total</sub>
39.753 (6)	99.935%E3 + 0.065%M4 $\delta = 0.025^{+10}_{-15}$	142 (7)	1051 (30)	13 (5)	423 (6)	615 (18)	214 (6)	31.4 (9)	0.024 (11)	1440 (40)

#### Recommended gamma-ray energy, absolute emission probability, and transition probability from 100% IT decay of <sup>103m</sup>Rh.

E <sub>γ</sub> (keV)	P <sub>γ</sub> <sup>abs</sup> (%)	Transition probability (%)	Δ (TP - P <sub>γ</sub> <sup>abs</sup> ) (%)
γ <sub>1,0</sub> 39.753 (6)	0.0694 (15)	100.0	99.9306 (15)

Electron energies were determined from electron binding energies tabulated by Larkins (1977La19) and the evaluated gamma-ray energies. Absolute electron emission probabilities were calculated from the evaluated absolute gamma-ray emission probabilities and associated internal conversion coefficients.

#### Energies and emission probabilities of internal conversion electrons from 100% IT decay of <sup>103m</sup>Rh.

		Energy (keV)	Electrons per 100 disint.
ec <sub>1,0 T</sub>	(Rh)	16.533 – 39.751	99.9306 (15)
ec <sub>1,0 K</sub>	(Rh)	16.533 (6)	9.85 (+50-40) → 9.85 (45)
ec <sub>1,0 L</sub>	(Rh)	36.341 – 36.749	73 (3)
ec <sub>1,0 M</sub>	(Rh)	39.126 – 39.446	14.9 (5)
ec <sub>1,0 N+</sub>	(Rh)	39.672 – 39.751	2.18 (8)

### Atomic Data

X-ray decay data and related parameters for <sup>103m</sup>Rh have been measured experimentally over many years (e.g., 1973In07, 1974Sa15, 1975Cz03, 1981Va22, 1994Sc43), and most recently by Riffaud *et al.* (2017Ri10, 2018Ri01, 2021Ri01). The <sup>103</sup>Rh(n,n')<sup>103m</sup>Rh reaction is important for reactor dosimetry studies because of the generation of suitable data traceable to reactor operations. This radionuclide emits only X-rays for which emission probabilities can be determined by means of the single gamma-ray transition probability and fluorescence yields. While these ω<sub>K</sub>, ω<sub>L</sub>, etc. have been measured as well as calculated by means of semi-empirical and theoretical models, available experimental data are rather limited as implied in the following table (1936Ba01, 1937St02, 1954Ro45, 1957Ro58, 1990Si19, 1992Ga24, 2017Ri10). Both the X-ray and Auger-electron data have been subsequently calculated on the basis of the evaluated gamma-ray data, along with atomic data from 1996Sc06, 1998ScZM, 1999ScZX, 2000Sc47 (Rh: ω<sub>K</sub> = 0.809 (4); ω<sub>L</sub> = 0.0494 (12); n<sub>KL</sub> = 0.987 (4)).

## Fluorescence yields for Rh.

Reference	$\omega_K$	$\omega_L$	$\omega_M$	Comments
1936Ba01	0.801 <sup>a</sup>	—	—	experimental measurements
1937St02	0.77 <sup>a</sup>	—	—	experimental measurements
1954Ro45	0.779 <sup>a</sup>	—	—	experimental measurements
1957Ro58	0.786 ± 0.015 <sup>b</sup>	—	—	experimental measurements
1990Si19/1994Hu23	0.829 ± 0.058	—	—	based upon 1990Si19 measurements of $\sigma_K$
1992Ga24/1994Hu23	—	0.051 ± 0.005	—	based upon 1992Ga24 measurements of $\omega_L$
2017Ri10	0.814 ± 0.041	—	—	experimental measurements
<b>Mean or only value</b>	<b>0.803 ± 0.024</b>	<b>0.050 ± 0.002<sup>c</sup></b>	<b>0.00135<sup>d</sup></b>	
1972Bb16	0.807 ± 0.031	—	—	assessment/evaluation of $\omega_K$ models
1994Hu23	0.792 0.8086	0.0499 0.0459	0.00135 0.00135	assessment/evaluation of $\omega_K$ , $\omega_L$ and $\omega_M$ models development of existing fitted functions
1996Sc06	0.809 ± 0.004	0.0494 ± 0.0012	—	assessment/evaluation of $\omega_K$ and $\omega_L$ models
2011Ka49	0.80205	—	—	assessment/evaluation of $\omega_K$ models

<sup>a</sup> Unspecified uncertainty – not included in weighted-mean analyses of experimental data set.<sup>b</sup> Uncertainty adjusted from ± 0.015 to ± 0.034 to reduce weighting to no more than 0.50.<sup>c</sup> Estimation, as guided by a single measurement (1992Ga24) and assessments/evaluations (1994Hu23, 1996Sc06).<sup>d</sup> Estimation, as proposed by assessment/evaluation (1994Hu23).Measured LX-ray decay data: <sup>103m</sup>Rh.

	Levels	Energy (keV)	Relative PLX
2021Ri01	L3-M1	2.370	2.67 (14)
	L2-M1	2.513	1.73 (9)
	L3-M4.5	2.695	100 (5)
	L2-M4	2.891	48.5 (24)
	L1-M2	2.891	8.23 (41)
	L1-M3	2.915	—
	L3-N1	2.934	—
	L3-N4.5	3.012	8.00 (40)
	L1-M4.5	3.101	4.94 (25)
	L2-N4	3.154	0.841 (43)
	L1-N2,3	3.373	0.848 (44)

X-ray and Auger-electron spectra for the IT decay of <sup>103m</sup>Rh were determined by means of BrIccEmis calculations within the NS\_RadList program, as described in 2012Le09, 2013LeZX, 2016Le19, 2020TeZY and 2023Texx. Transition energies within each propagation step were derived from the atomic binding energies determined by adopting the relativistic Dirac-Fock approach employed in the RAINe code (2002Ba85), along with the application of a semi-empirical correction procedure that aligns these energies more fully with known spectral data. Fixed transition rates were obtained from the EADL database (1991PeZY, 1993Cu08). Resulting X-ray and Auger-electron energies and emission probabilities are listed below for <sup>103m</sup>Rh probabilities (both sets of energies are listed as mean energies and ranges of energies at the 95% confidence level, with additional extensions of these energy ranges in italics for specific X-ray and Auger electron emissions). The number of simulated nuclear events is 10<sup>6</sup>, and the intensity cut-offs for both the X-ray and Auger-electron listings are 0.01 per 100 decays. As defined by 1991PeZY, uncertainties in these theoretical X-ray emission probabilities are ~ 10% for the K and L shells and 30% for the outer shells, whereas uncertainties in the theoretical Auger-electron emission probabilities are < 15% for the K and L shells (except for Coster-Kronig and super Coster-Kronig transitions) and 30% for the outer shells.

X-ray energies and emission probabilities from IT decay of <sup>103m</sup>Rh (BrIccEmis/NS\_RadList).

	Mean energy (keV)	Energy (keV) 95% confidence range/extension	Photons per 100 decays <sup>a</sup>
X <sub>tot</sub> (Rh)	14.255	2.378 – 22.725	12.5 (4)
XK <sub>tot</sub> (Rh)	20.599	20.073 – 23.167/23.173	8.1 (4)
XKL2 (Rh)	20.073	20.073	2.33 (10)
XKL3 (Rh)	20.215	20.215	4.41 (20)
XKM (Rh)	22.717	22.699 – 22.725	1.11 (5)
XKM2 (Rh)	22.699	22.699	0.375 (17)
XKM3 (Rh)	22.725	22.725	0.730 (33)
XKN (Rh)	23.171	23.167 – 23.173	0.219 (10)
XKN2 (Rh)	23.167	23.167	0.0746 (34)
XKN3 (Rh)	23.173	23.173	0.144 (6)
XL <sub>tot</sub> (Rh)	2.743	2.378 – 3.067	4.27 (10)
XM <sub>tot</sub> (Rh)	0.321	0.189 – 0.546	0.1184 (23)

<sup>a</sup> Quoted uncertainties include direct contributions from only the internal-conversion electrons (2023Texx) – these data do not include significant uncertainties from the more complex but less well-defined atomic radiation probabilities as considered semi-quantitatively in EADL, whereby uncertainties in the theoretical X-ray radiative rates are ~ 10% for the K and L shells, and can be of the order of 30% for the outer shells (1991PeZY).

Auger-electron energies and emission probabilities from IT decay of <sup>103m</sup>Rh (BrIccEmis/NS\_RadList).

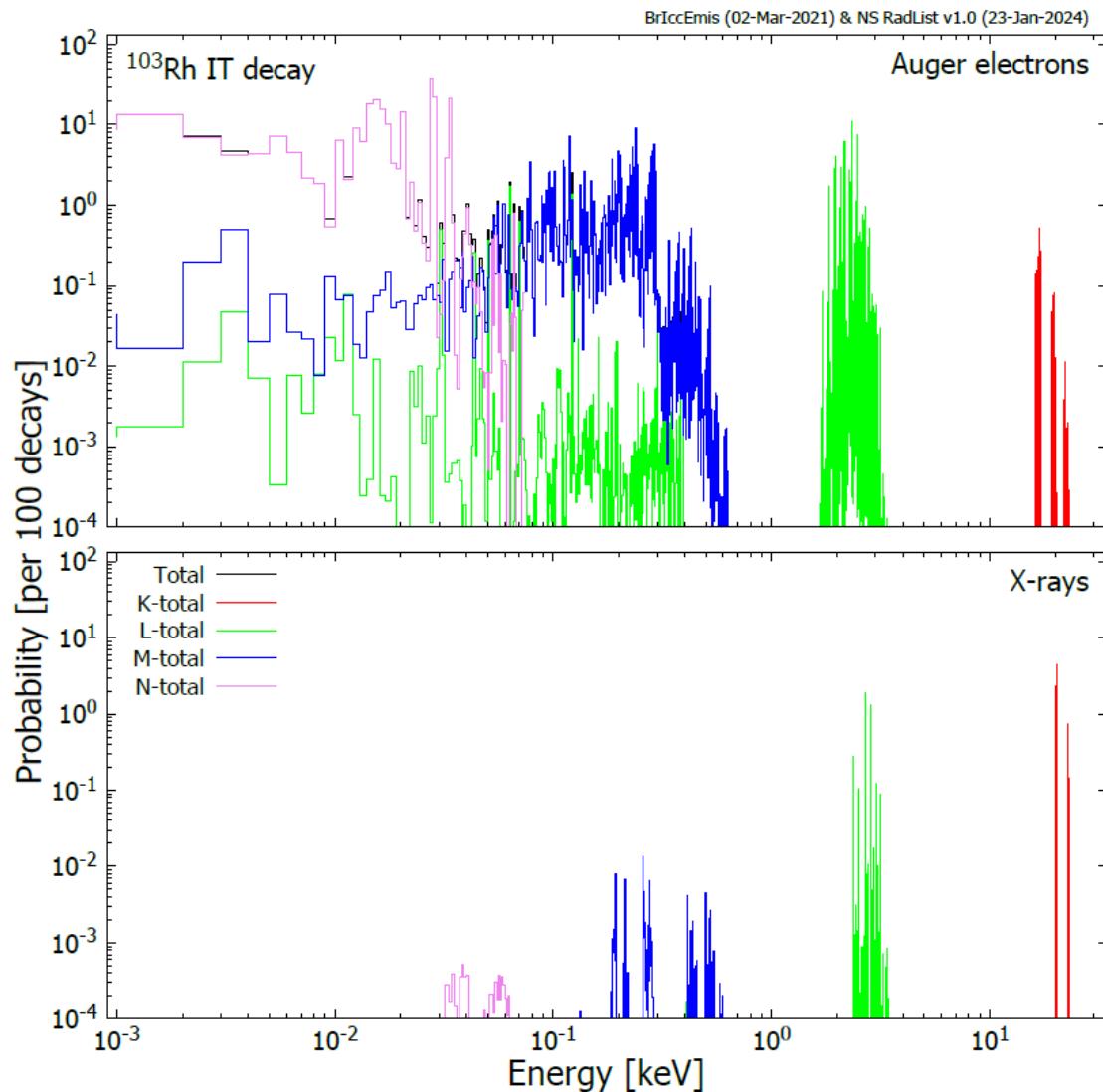
		Mean Energy (keV)	Energy (keV) 95% confidence range/extension	Electrons per 100 decays <sup>a</sup>
Auger total	(Rh)	0.461	0.002 – 2.490	582 (12)
Auger K <sub>tot</sub>	(Rh)	17.708	16.281 – 21.922/22.838	1.92 (9)
Auger KLL	(Rh)	16.811	16.281 – 17.083	1.34 (6)
Auger KLX	(Rh)	19.574	19.133 – 20.146	0.525 (24)
Auger KXY	(Rh)	22.275	21.922 – 22.838	0.0495 (22)
Auger L <sub>tot</sub>	(Rh)	2.139	0.004/0.064 – 2.768/3.133	85.4 (19)
Auger Coster-Kronig LLM	(Rh)	0.038	0.004 – 0.051	0.94 (+22-17)
Auger Coster-Kronig LLX	(Rh)	0.098	0.031 – 0.313	5.36 (15)
Auger LMM	(Rh)	2.239	1.842 – 2.490	66.2 (15)
Auger LMX	(Rh)	2.600	2.383 – 2.824	12.25 (28)
Auger LXY	(Rh)	2.969	2.851 – 3.133	0.620 (14)
Auger M <sub>tot</sub>	(Rh)	0.200	0.021/0.054 – 0.368/0.390	235 (5)
Auger Coster-Kronig MMX	(Rh)	0.103	0.021 – 0.178	69.0 (13)
Auger MXY	(Rh)	0.240	0.141 – 0.390	166.1 (34)
Auger N <sub>tot</sub>	(Rh)	0.019	0.001 – 0.034/0.064	259 (5)
Auger super Coster-Kronig NNN	(Rh)	0.019	0.001 – 0.034	249 (5)
Auger Coster-Kronig NNX	(Rh)	0.020	0.002 – 0.064	9.94 (21)

<sup>a</sup> Quoted uncertainties include direct contributions from only the internal-conversion electrons (2023Texx) – these data do not include significant uncertainties from the more complex but less well-defined atomic radiation probabilities as considered semi-quantitatively in EADL, whereby uncertainties in the theoretical Auger-electron radiative rates are < 15% for the K and L shells (except for Coster-Kronig and super Coster-Kronig transitions), and can be of the order of 30% or more for Coster-Kronig and super Coster-Kronig transitions and the outer shells (1991PeZY).

Total energy release per decay as determined from the BrIccEmis/NS\_RadList studies:

Radiation	Total energy (keV per decay)	Total intensity (%)
Gamma rays	0.0276 (8)	0.0694 (19)
Conversion electrons	35.3 (12)	100.3 (35)
Auger electrons	2.67 (6)	582 (12)
X-rays	1.77 (6)	12.5 (4)
Summed total	39.8 (12)	
Q(IT)	39.753 (6)	

The final X-ray and Auger-electron spectra were effectively evaluated from one million simulated nuclear decay events. Both component spectra of the Auger-electron and X-ray emissions as calculated by the BrIccEmis code for the IT decay of <sup>103m</sup>Rh are shown in the following figure.

Auger-electron (upper panel) and X-ray (lower panel) spectra of <sup>103m</sup>Rh IT decay:Main Production Modes for <sup>103</sup>Pd and <sup>103m</sup>Rh

<sup>103m</sup>Rh:  $^{103}\text{Rh}(\text{n},\text{n}')^{103m}\text{Rh}$ ,  $^{103}\text{Rh}(\text{d},\text{d}')^{103m}\text{Rh}$ ,  $^{103}\text{Rh}(\alpha,\alpha')^{103m}\text{Rh}$ ,  $^{103}\text{Rh}(\gamma,\gamma')^{103m}\text{Rh}$   
 parent  $^{103}\text{Pd}(\text{EC})^{103m}\text{Rh}$   
 parent  $^{103}\text{Ru}(\beta)^{103m}\text{Rh}$ :  $^{102}\text{Ru}(\text{n},\gamma)^{103}\text{Ru}$ ,  $^{102}\text{Ru}(\text{d},\text{p})^{103}\text{Ru}$ ,  $^{104}\text{Ru}(\text{d},\text{t})^{103}\text{Ru}$ ,  $^{100}\text{Mo}(\alpha,\text{n})^{103}\text{Ru}$

## Data Consistency

<sup>103m</sup>Rh

A Q<sub>IT</sub>-value of 39.753 (6) keV for the IT decay of <sup>103m</sup>Rh from NUBASE2020/AME2020 (2021Ko07, 2021Wa16) has been similarly compared with a calculated Q-value of 39.8 (6) keV to give a percentage deviation of  $-(0.1 \pm 1.5)\%$ , which supports the derivation of a consistent decay scheme with a modest variant.

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The above codes have undergone and will continue to undergo various forms of modification and improvement, which will be reflected in changes to the version number and date of issue into the public domain. Assistance from Tibor Kibédi (ANU, Australia – particularly BrIccEmis and NS\_RadList) concerning improved operational features of some of the above codes is gratefully acknowledged by ALN.